Reionization and the cosmic dawn

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Epoch of Reionization

Radiative Transfer Code – C^2 -Ray

- working principle
- current problems

New Algorithm

- Adaptive timestep scheme
- Asynchronous evolution

Conclusion

Epoch of reionization



Reionization simulation





N-body simulation

Radiative transfer simulation

Radiative Transfer equation

The complete RT equation

$$\frac{1}{c}\frac{\partial I_{\nu}}{\partial t} + \frac{\hat{n}\cdot\nabla I_{\nu}}{\bar{a}} - \frac{H(t)}{c}\left(\nu\frac{dI_{\nu}}{d\nu} - 3I_{\nu}\right) = j_{\nu}(\mathbf{x},\hat{n}) - \alpha_{\nu}(\mathbf{x},\hat{n})I_{\nu}$$

In numerical simulation, we approximate it to...

$$\hat{n} \cdot \nabla I_{\nu} = j_{\nu}(\mathbf{x}, \hat{n}) - \alpha_{\nu}(\mathbf{x}, \hat{n})I_{\nu}$$

But still it is an expensive calculation since it is a 7dimensional problem

Radiative Transfer code - C²-Ray

Ist version (Mellema, 2006) – grid-based, photon conserving, hydrogen only, short characteristic,
 On the spot approximation, iteration of the photo equations using averaged results

2nd version (Friedrich, 2012) – helium inclusion, multifrequency approach, secondary ionization

3rd version (Lee, in preparation) –



We solve the ionization and thermal equations

$$rac{dx}{dt} = (1-x)(\Gamma + n_e C_H) - x n_e lpha_H \qquad \qquad rac{\partial u}{\partial t} = \mathcal{H} - \mathcal{C}$$

The solutions are only accurate if the

- I. Photoionization rate (Γ) and
- 2. Photoheating rate (H) are estimated correctly

$$\Gamma_{HI}(\tau_{\nu}) = \int_{\nu_{HI}}^{\infty} \frac{L_{\nu}}{h\nu} \frac{e^{-\tau_{\nu}} (1 - e^{-\tau_{\nu}})}{n_{HI}} \frac{1}{V_{shell}} d\nu$$

$$\mathcal{H}_{HI}(\tau_{\nu}) = \int_{\nu_{HI}}^{\infty} \frac{L_{\nu}}{h\nu} e^{-\tau_{\nu}} (1 - e^{-\tau_{\nu}}) \frac{h(\nu - \nu_{HI})}{V_{shell}} d\nu$$

Problem of C²-Ray

C2-Ray uses the below Γ and H

$$\Gamma_{HI}(\langle \tau_{\nu} \rangle) = \int_{\nu_{HI}}^{\infty} \frac{L_{\nu}}{h\nu} \frac{e^{-\langle \tau_{\nu} \rangle} (1 - e^{-\langle \tau_{\nu} \rangle})}{\langle n_{HI} \rangle} \frac{1}{V_{shell}} d\nu$$
$$\mathcal{H}_{HI}(\langle \tau_{\nu} \rangle) = \int_{\nu_{HI}}^{\infty} \frac{L_{\nu}}{h\nu} e^{-\langle \tau_{\nu} \rangle} (1 - e^{-\langle \tau_{\nu} \rangle}) \frac{h(\nu - \nu_{HI})}{V_{shell}} d\nu$$

The correct $\[\]$ and $\[\]$ are

$$\langle \Gamma_{HI}(\tau_{\nu}) \rangle = \int_{\nu_{HI}}^{\infty} \frac{L_{\nu}}{h\nu} \left\langle \frac{e^{-\tau_{\nu}} (1 - e^{-\tau_{\nu}})}{n_{HI}} \right\rangle \frac{1}{V_{shell}} d\nu$$

$$\langle \mathcal{H}_{HI}(\tau_{\nu}) \rangle = \int_{\nu_{HI}}^{\infty} \frac{L_{\nu}}{h\nu} \left\langle e^{-\tau_{\nu}} (1 - e^{-\tau_{\nu}}) \right\rangle \frac{h(\nu - \nu_{HI})}{V_{shell}} d\nu$$

However, C2-Ray uses long time-step, it leads to $\langle \Gamma_{HI}(\tau_{\nu}) \rangle \approx \Gamma_{HI}(\langle \tau_{\nu} \rangle) \qquad \langle \mathcal{H}_{HI}(\tau_{\nu}) \rangle \gg \mathcal{H}_{HI}(\langle \tau_{\nu} \rangle)$

Problem of C²-Ray

lonization evolution : long time-step is ok.

Thermal evolution : long time-step is problematic.





We need to use a shorter time-step

Traditional time-step

Time-step depends on timescale of some physical process

- I. Ionization rate
- 2. Heating
- 3. Hydrodynamic

Scan through entire grid, choose the minimum one.

- Our adaptive time-step bases on one principle: Ionization front does not move more than one cell after N time-step
- N is a parameter determined upon the time of calculation
- It gives more control over the evolution
- We need to identify the ionization front cells that can be reached by each source
- Since we do this procedure in every time-step, we must contrive some quick method





Domain decomposition I cube \rightarrow 24 pyramids



Take one pyramid as an example, the green cells are ionization front, are they all reached by the yellow source?



We can send many rays out and check the first green cells met by each ray, BUT....



Our algorithm !!

We send the first ray and identify the ionization front cell met by this ray



We send the second ray so that it passes just above the previous met cell





and so on



and so on



and so on



and so on

The first slice is complete, we move to the second slice



and so on



and so on



and so on



and so on



and so on



and so on



and so on



and so on



and so on



and so on



and so on



and so on

and so on (This ray does not meet ionization front cell)





and so on



and so on



and so on



and so on



and so on



and so on



and so on



and so on



and so on



and so on



and so on





and so on

Trick I – Adaptive timestep

Finish!

These are all the rays that we sent, they are less numerous than the previous one



In an extreme case, the source generates a small ionized bubble while other sources generate bigger ionized bubble. However, we use only 4 rays to complete the problem

- Given heating timescale << recombination timescale
- Time-step ~ heating timescale is essential to neutral cells
- Time-step ~ recombination timescale is enough for ionized cells
- Applying short time-step to ionized cells is a waste
- Asynchronous evolution : applying long time-step to ionized cells and short time-step to neutral cells
- At the intermediate stage, the clocks of ionized cells and neutral cells are not synchronous, but will finally be synchronized

Given a grid of gas and 5 sources.

I want to see what happens after Δt First, the grid evolves with a short time-step Δt_1





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y-z plane

x-z plane

x-y plane

The pink cells are ionized Next, I apply short time-step Δt_2 to the neutral cells and $\Delta t - \Delta t_1$ to the ionized cell





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y-z plane

x-z plane

x-y plane

More cells are ionized Next, I apply short time-step Δt_3 to the neutral cells and $\Delta t - \Delta t_2$ to the newly ionized cell





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y-z plane

x-z plane

x-y plane

More cells are ionized Next, I apply short time-step Δt_4 to the neutral cells and $\Delta t - \Delta t_3$ to the newly ionized cell





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y-z plane

x-z plane

x-y plane

More cells are ionized Next, I apply short time-step Δt_5 to the neutral cells and $\Delta t - \Delta t_4$ to the newly ionized cell





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y-z plane

x-z plane

x-y plane

Now Δt has passed. By allowing cells evolve with different time-step, the actual number of evolutions is reduced. Computational cost is also reduced!





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y-z plane

x-z plane

x-y plane

Summary

To accurately calculate thermal evolution, short time-step is required.

Finding correct time-step can be expensive Solved by strategy - Adaptive time-step

Evolution of entire domain with short time-step is very expensive

Solved by strategy - Asynchronous evolution